

Coverage of Error Bars for Poisson Data

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Abstract

The frequentist concept of coverage is explained, and illustrated by calculating the coverage properties of eight error bar schemes for Poisson data. The primary goal is to aid physicists in doing their own coverage calculations, but some conclusions are also drawn concerning the relative coverage performance of the eight schemes. While mostly intended for beginners, some advanced concepts are also introduced.

1 Introduction

When physicists determine an unknown parameter from experimental data, they also provide error bars. The central value and the error bars, often written $V_{-\sigma_1}^{+\sigma_2}$, determine an interval $[V - \sigma_1, V + \sigma_2]$, which is the region within the error bars. One quantity of interest associated with such an interval is the coverage probability (usually just called the coverage) of the interval. We define coverage in the following single-parameter example:

We will assume that there is a single unknown parameter μ that is estimated from the data \vec{x} . The experimenters have functions that determine the central value and errors from the data— $V(\vec{x})$, $\sigma_1(\vec{x})$, and $\sigma_2(\vec{x})$ respectively—for all possible data \vec{x} . (These functions often are defined by, and implemented through, some fitting procedure.)

The data follow a probability distribution $p(\vec{x}, \mu)$ that depends on μ , and is completely known once a specific numerical value for μ is picked. The coverage $C(\mu)$ is a function of μ , defined as the probability that

$$V(\vec{x}) - \sigma_1(\vec{x}) \leq \mu \leq V(\vec{x}) + \sigma_2(\vec{x})$$

for random \vec{x} generated from $p(\vec{x}, \mu)$. It is important to note that in this equation, μ is regarded as fixed, and the probability statement applies to the “variables” (really functions) V , σ_1 , and σ_2 , which in turn depend on \vec{x} .

So the coverage is a function $C(\mu)$ of the unknown parameter. It is, for any given μ , the probability that an experiment, with data following the same distribution $p(\vec{x}, \mu)$, and employing the same analysis functions $V(\vec{x})$, $\sigma_1(\vec{x})$, and $\sigma_2(\vec{x})$, will obtain an interval that includes (i.e. “covers”) μ . The Holy Grail of Frequentist Statistics is to define $V(\vec{x})$, $\sigma_1(\vec{x})$, and $\sigma_2(\vec{x})$ so that $C(\mu) = C_0$, C_0 being some predefined constant. This optimum case (constant $C(\mu)$) is referred to as exact coverage. By default, physicists use, and will normally assume, $C_0 = 0.682689492137\dots$ (the area of the Gaussian distribution contained within $\pm 1\sigma$) as the coverage value for error bars¹: if any other coverage value is used (for error bars), it must be explicitly stated to avoid misunderstandings.

Exact coverage, like the Grail of legend, if approached by any but a perfectly pure and holy frequentist, is borne away and vanishes from sight. So, the following examples mainly demonstrate how badly the coverage deviates from exact when applying common methods to Poisson distributed data.

In the sections that follow we will investigate the coverage $C(\mu)$ achieved by eight different error-bar schemes for Poisson data with no background. In fact, exact coverage cannot be achieved through normal means for the Poisson distribution, or discrete distributions in general—the reasons for this will be manifest upon understanding the examples.

2 Pearson’s χ^2 Intervals

In this section we examine the coverage of intervals derived from Pearson’s χ^2 . [1, 2] Specifically, we choose a Poisson process characterized by parameter $\mu \geq 0$ from which we observe n events. The probability of observing n events is

$$p(n, \mu) = \frac{e^{-\mu} \mu^n}{n!}$$

and Pearson’s χ^2 is given by

$$\chi^2(\mu, n) = \frac{(n - \mu)^2}{\mu}$$

(We give the arguments of the probability as (n, μ) to indicate that n is variable and μ is fixed. For the χ^2 , it’s μ that is variable and n that is fixed, so the arguments are swapped.)

¹There are other physics conventions for limits, where $C(\mu) = 95\%$ is a common choice.

Having observed n events, we obtain our central estimate V of the unknown parameter μ by minimizing the χ^2 with respect to μ , obtaining $V = n$. For the error bars, we adopt the interval defined as the set of all μ such that $\chi^2(\mu, n) \leq \Delta$. In the language of Minuit[3], these are MINOS errors, and Δ is the Minuit `ERRDEF` parameter. The MINOS error bars are defined as the change in parameter value (μ) required to increase the function value (χ^2) by `ERRDEF` (Δ). In this simple example, the errors are

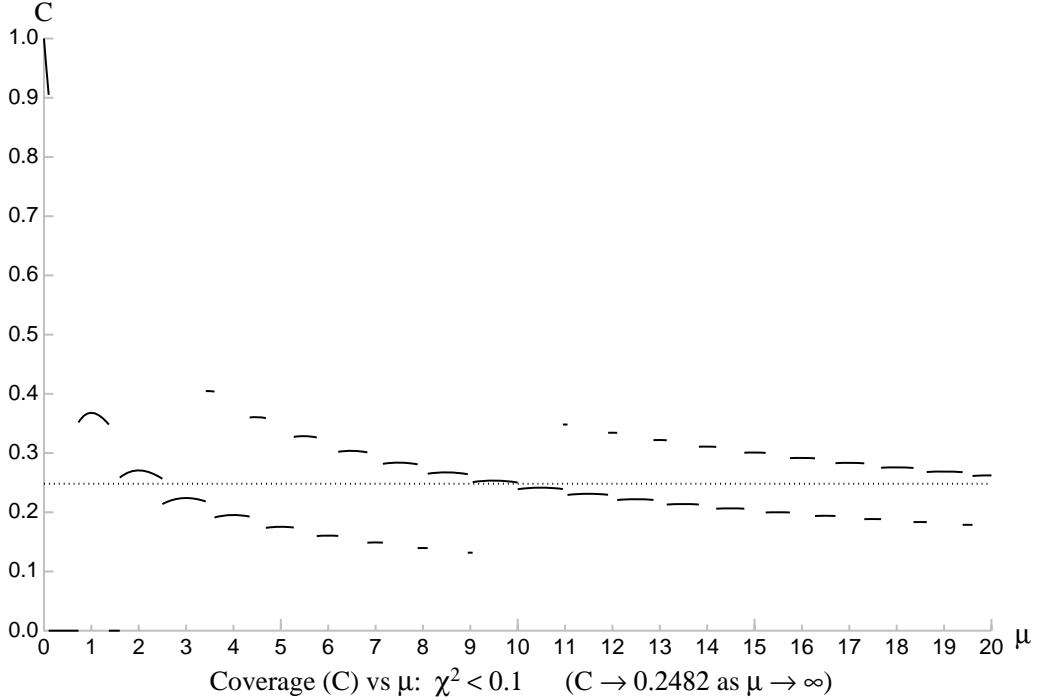
$$\sigma_1 = \sqrt{n\Delta + \Delta^2/4} - \Delta/2 \quad \sigma_2 = \sqrt{n\Delta + \Delta^2/4} + \Delta/2$$

and we also have the useful relations $\sigma_1\sigma_2 = n\Delta$ and $\sigma_2 - \sigma_1 = \Delta$. For no observed events ($n = 0$), the above formulas still are valid, and we have $\sigma_1 = 0$ and $\sigma_2 = \Delta$.

The interval $[\mu_1, \mu_2]$ that defines these error bars satisfies

$$\mu_1 = n + \Delta/2 - \sqrt{n\Delta + \Delta^2/4} \quad \mu_2 = n + \Delta/2 + \sqrt{n\Delta + \Delta^2/4}$$

and $\mu_1\mu_2 = n^2$, and we wish to calculate the coverage of this interval. Note that the size of the interval, $\sqrt{4n\Delta + \Delta^2}$, grows steadily with the observed number of events n . At this point in the discussion, it is useful to show a plot of the coverage $C(\mu)$ for the case $\Delta = 0.1$:



$\Delta = 0.1$ is not the usual choice; this plot is shown first because it is easier to explain than the plot to follow with $\Delta = 1$. There are two shocks to recover from: $C(\mu)$ is discontinuous at ~ 40 points in the region $\mu < 20$, and the value of $C(\mu)$ is all over the map. The first time one sees a plot like this, one assumes that some bug must exist, but, of course, the plot turns out to be correct.

The following explanation should help: At $\mu = 0$, the coverage must be 100%—zero events are always observed, and the $n = 0$ interval $[0, 0.1]$ contains $\mu = 0$ every time. As μ increases slightly, but remains less than 0.1, occasionally $n \geq 1$ is observed, but the interval for $n = 1$ is $[0.7298, 1.3702]$, which does not cover $\mu < 0.1$. So $C(\mu) = p(0, \mu) = e^{-\mu}$ for $\mu < 0.1$. For μ greater than 0.1 and less than 0.7298, no possible n has an error interval that covers μ . So $C(\mu) = 0$ for $0.1 < \mu < 0.7298$. When μ is within the region $[0.7298, 1.3702]$, only $n = 1$ covers, since the $n = 2$ interval is $[1.6, 2.5]$ (exactly). So $C(\mu) = p(1, \mu) = \mu e^{-\mu}$ for $\mu \in [0.7298, 1.3702]$.

Similarly, after another zero coverage region, we have $C(\mu) = p(2, \mu) = \mu^2 e^{-\mu}/2$ for $\mu \in [1.6, 2.5]$. The interval when $n = 3$ is observed is $[2.5, 3.6]$ (exactly), and the interval for $n = 4$ is $[3.4156, 4.6844]$, which overlaps the $n = 3$ interval. So, for $\mu \in [2.5, 3.4156]$, $C(\mu) = p(3, \mu) = \mu^3 e^{-\mu}/6$. For $\mu \in [3.4156, 3.6]$, both the $n = 3$ and $n = 4$ cases cover, and $C(\mu) = p(3, \mu) + p(4, \mu) = \mu^3 e^{-\mu}/6 + \mu^4 e^{-\mu}/24$.

After that, the form $C(\mu) = (k, \mu)$ alternates with $C(\mu) = p(k, \mu) + p(k + 1, \mu)$ for a while. For $\mu > 10$, the error intervals are wide enough so that there are regions where three of them overlap, and $C(\mu) = p(k, \mu) + p(k + 1, \mu)$ alternates with $C(\mu) = p(k, \mu) + p(k + 1, \mu) + p(k + 2, \mu)$.

Since discontinuities in $C(\mu)$ occur at the beginning and the end of each interval, within the region $\mu < 20$ there are about 40 discontinuities. This property—average continuous segment width of 0.5—must hold quite accurately over large regions. Strictly speaking, there are two segments of zero width not mentioned above: for $\mu = 2.5$ exactly, under our definition, both the $n = 2$ and $n = 3$ intervals cover, so $C(2.5) = p(2, 2.5) + p(3, 2.5) = 0.4703$. But line segments of zero size simply don't show up on the plot. The other orphan coverage point is at $\mu = 10$.

The following trivial C program suffices to calculate the coverage of Pearson's χ^2 intervals for any μ small enough that $e^{-\mu}$ does not underflow. It expects two command-line arguments: the first is μ , and the second is Δ . As the entire calculation only takes a half a dozen lines or so of C-code, a careful examination of the logic should reward the reader with an improved understanding.

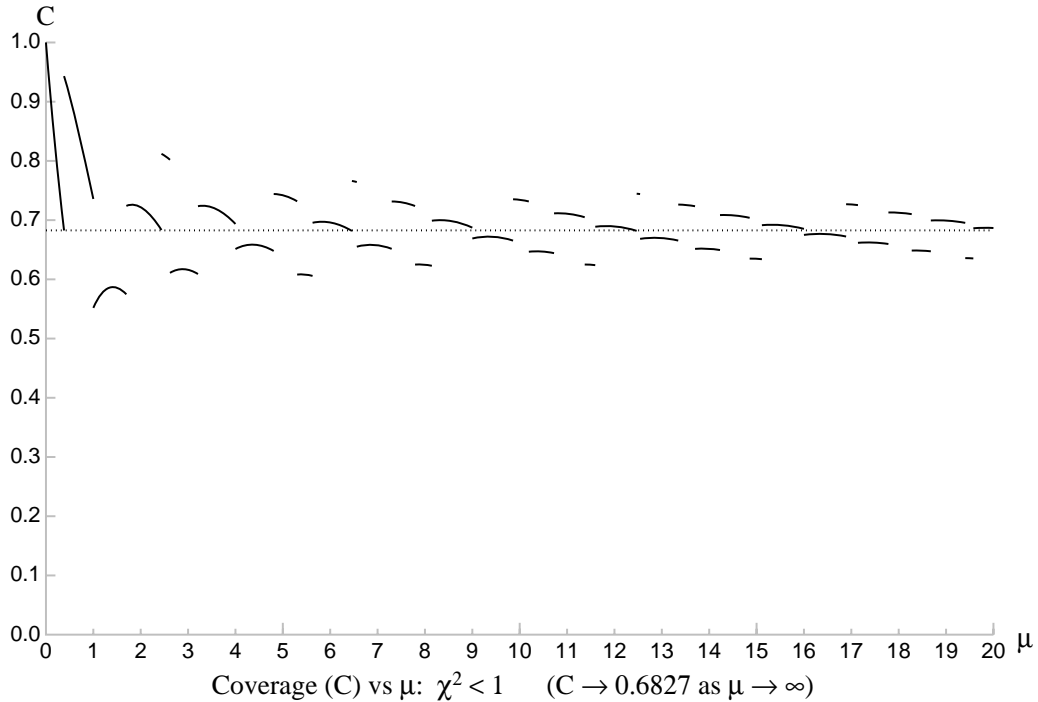
```

#include <stdio.h>
#include <stdlib.h>
#include <math.h>
double chi2(double mu,int n) { return (n) ? (n-mu)*(n-mu)/mu : mu; }

int main(int argc, char* argv[]) {
    const double mu = (argc>1) ? strtod(argv[1],NULL) : 0.0;
    const double delta = (argc>2) ? strtod(argv[2],NULL) : 1.0;
    double sum=0.0, p=exp(-mu);
    int n;
    for(n=0 ; p>0 ; p *= mu/(++n)) {
        if( chi2(mu,n) <= delta )
            sum += p;
        else if(n>mu)
            break;
    }
    printf("mu=%g delta=%g coverage=%g\n",mu,delta,sum);
    return 0;
}

```

The following plot shows $C(\mu)$ for $\Delta = 1$:



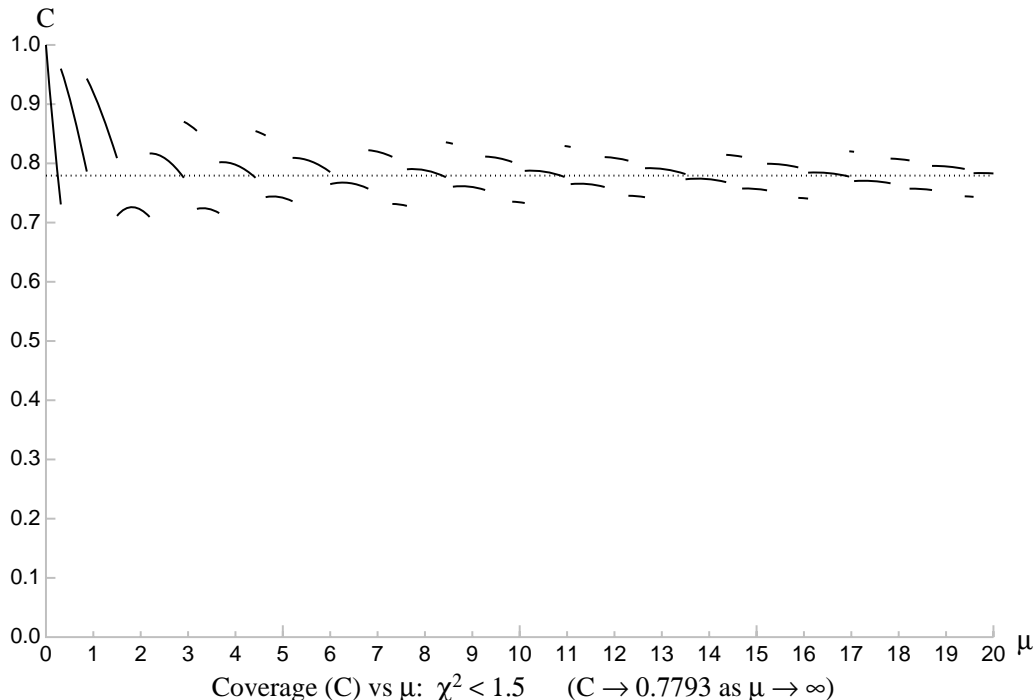
$\Delta = 1$, of course, is the physicist's standard choice for 1σ error bars. The minimum value for $C(\mu)$ on this plot is $1.5e^{-1} = 0.5518$, which is attained in

the limit as μ approaches 1 from above. There is an orphan point at exactly $\mu = 1$, and $C(1) = 2.5e^{-1} = 0.9197$. (To the attentive reader, it will be obvious from looking at the plot that there are also orphan points at $n = 4$, $n = 9$, and $n = 16$.) It seems amazing that one obtains such a complicated structure from such a simple rule. The program given above is sufficient to calculate the coverage for a given value of μ , but it won't indicate where the discontinuities are located; the discontinuities occur at the interval boundaries. The plot is actually produced by locating all the discontinuities first, sorting them in ascending order, and then plotting $C(\mu)$ as continuous curves between the discontinuities.

Suppose we do an experiment and we observe 6 events. Under the above rule, we report $\mu = 6^{+3}_{-2}$. From the frequentist point of view, μ is still an unknown parameter: it could be in the neighborhood of 1 (where the minimum coverage occurs). All we can say about the coverage with absolute certainty is that it is greater than or equal to 55.18% (and less than or equal to 1). Average coverage is a Bayesian concept; for the frequentist, μ , although unknown, has a definite and fixed value². In fact, even if 100 events were observed, the $C \geq 55.18\%$ conclusion is still the only strictly valid statement, although in practice most frequentists would grant that $C \simeq 68.27\%$ is a reasonable approximation for 100 observed events.

But frequentists will generally demand (if the number of observed events is not large) that the minimum coverage be at least 68.27%. One simple strategy is just to boost Δ until this is achieved. In fact, this is first achieved when $\Delta = 1.5$ (minimum coverage as a function of Δ is also discontinuous). This case is shown in the following plot:

²The Bayesian and the frequentist both agree that μ has a definite and fixed value; the Bayesian assigns a probability distribution to μ to represent the prior state of his knowledge of what that true value might be.



Here the minimum coverage is 0.7095, achieved in the neighborhood of $\mu = 2.1883$. Unfortunately this leads to serious overcoverage³: the average coverage is $\sim 78\%$. When physicists see a plot that overcovers significantly (i.e., the error bars cover the theoretical curve at greater than 68%), they tend to accuse the authors of overstating their errors—or of being biased by the theory[4].

3 Neyman's Modified χ'^2 Intervals

Instead of Pearson's χ^2 , often Neyman's modified[5] χ'^2 is used:

$$\chi'^2(\mu, n) = \frac{(n - \mu)^2}{n}$$

where the number of observed events n replaces μ in the denominator. Although asymptotically approaching Pearson's χ^2 for large n , χ'^2 is generally thought to be inferior to Pearson's χ^2 at small n . One must also make some choice about what to do when $n = 0$. Once again, we take the error interval

³It also overcovers for all μ , since $C(\mu)$ is greater than 68.27% everywhere.

to be the set of all μ such that $\chi'^2(\mu, n) \leq \Delta$. The errors, symmetric for $n \geq \Delta$, are then given by

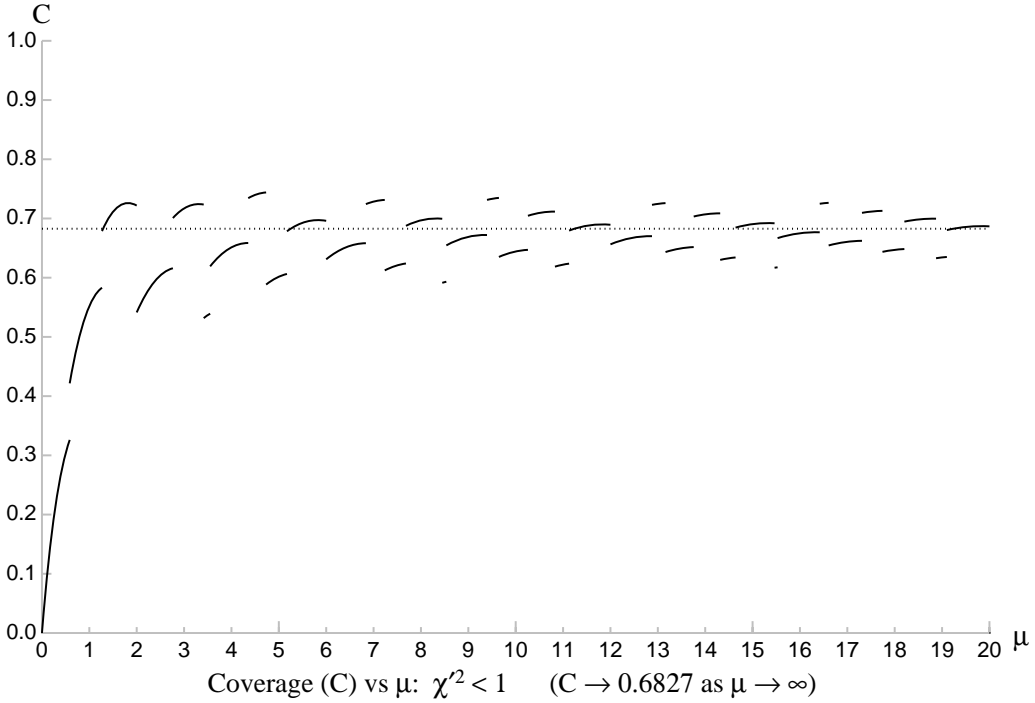
$$\sigma_1 = \min(n, \sqrt{n\Delta}) \quad \sigma_2 = \sqrt{n\Delta}$$

the interval is defined by

$$\mu_1 = \max(n - \sqrt{n\Delta}, 0) \quad \mu_2 = n + \sqrt{n\Delta}$$

and we have $\mu_1\mu_2 = n \max(n - \Delta, 0)$. For $n = 0$, we take $\sigma_1 = \sigma_2 = 0$, which is simply the limit of the above expressions.

The coverage $C(\mu)$ for the standard case $\Delta = 1$ is shown here:



For $0 < \mu < 2 - \sqrt{2}$, only the interval corresponding to $n = 1$ includes μ , so $C(\mu) = \mu e^{-\mu}$ in this region, and the coverage approaches zero in the limit as $\mu \rightarrow 0$. There is an orphan coverage point at $\mu = 0$: $C(0) = 1$ by our definition, since only in that special case is μ included in the $n = 0$ interval. There are also orphan coverage points at $\mu = 2, 6, 12$, and 20 .

From the frequentist point of view, the fact that the minimum coverage is zero (for any choice of Δ) is the worst possible outcome. This fate could have been avoided by picking some other (ad hoc) interval for $n = 0$, but we thought it best instead to illustrate what goes wrong in the given case.

Comparing with the corresponding Pearson's χ^2 plot on page 5, it is interesting that there the continuous coverage segments have a general negative slope, while here the segments, although arranged in a similar pattern, have a predominantly positive slope. Neither trend seems present in the $-2 \ln \lambda$ plot of page 10 in the next section, where the continuous segments tend to center more closely about their peak location.

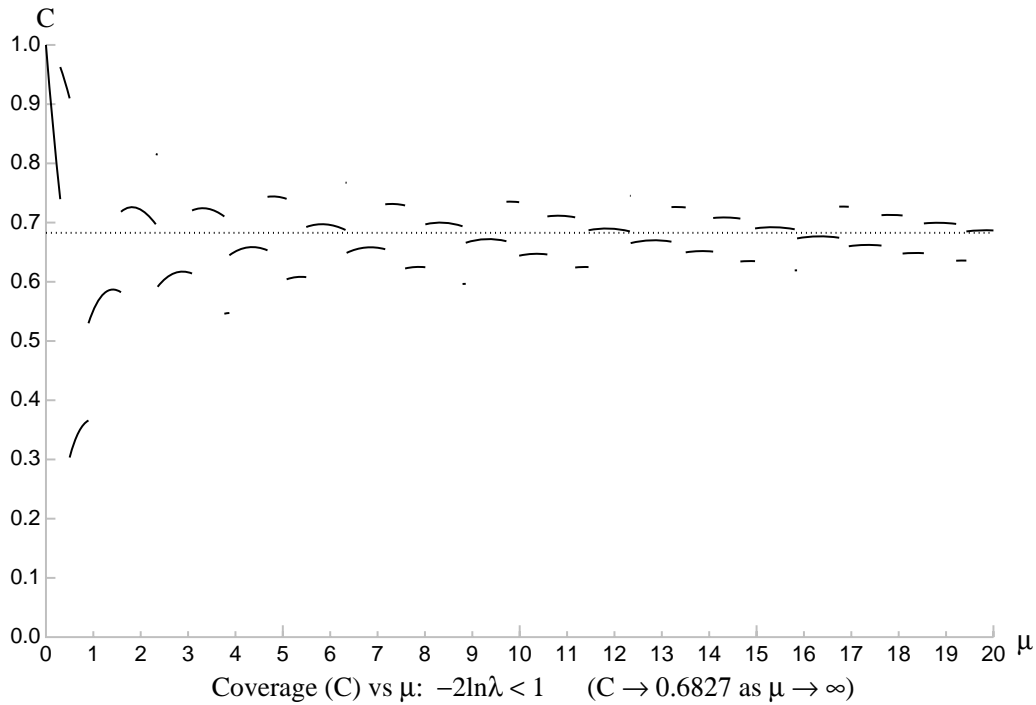
4 Likelihood Intervals

Instead of Pearson's χ^2 , or Neyman's modified χ'^2 , we can also try error intervals based on the value of the likelihood. Specifically, having observed n events, we can use the error interval defined as the set of all μ such that $-2 \ln \lambda(\mu, n) \leq \Delta$, where

$$-2 \ln \lambda(\mu, n) = 2[(\mu - n) + n \ln(n/\mu)]$$

is -2 times the log likelihood ratio⁴ of the Poisson distribution[6, 7]. This is the quantity that is minimized when one does a maximum-likelihood fit to the Poisson distribution. Once again, standard error bars correspond to $\Delta = 1$. The next plot shows the coverage for this case:

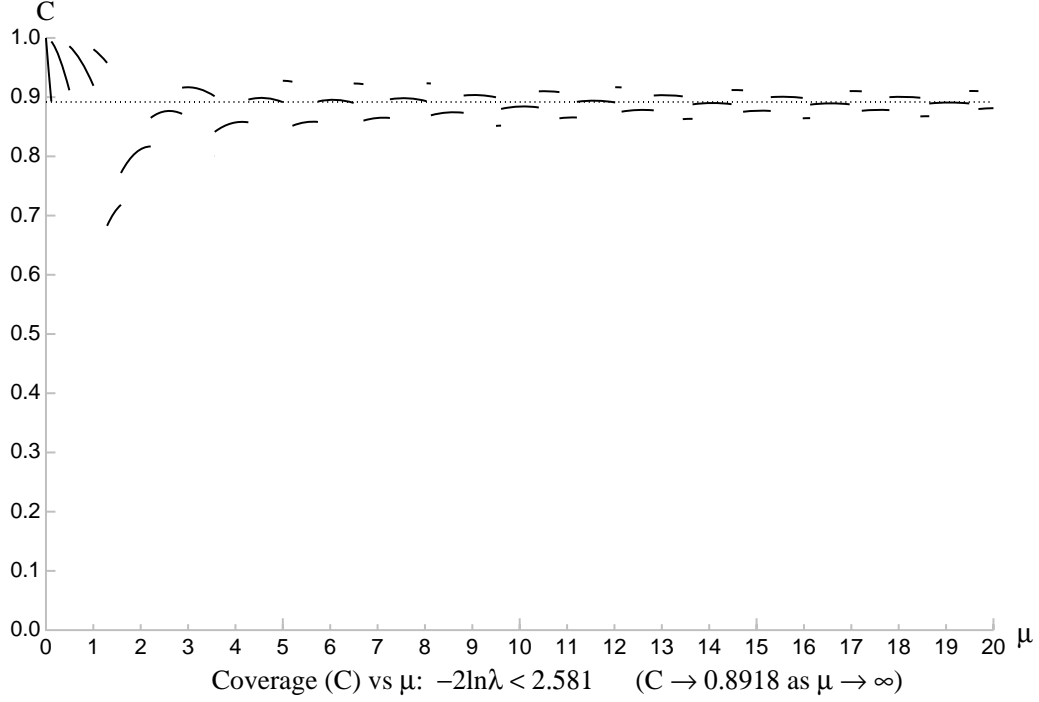
⁴The likelihood ratio $\lambda = p(n, \mu)/p(n, \mu_{\text{best}})$, where μ_{best} is the value of μ that maximizes $p(n, \mu)$ (n being treated as constant). Considered as a function of μ , λ is simply the likelihood renormalized so that the maximum value it can take is 1. In the Poisson case, $\lambda = p(n, \mu)/p(n, n)$. By definition, maximizing the likelihood (with respect to μ) is equivalent to maximizing λ or minimizing $-2 \ln \lambda$.



The minimum coverage, 0.3033, occurs in the neighborhood of $\mu = 0.5$. Surprisingly⁵, this is worse than for $\Delta = 1$ Pearson's χ^2 intervals. There are no orphan coverage points—those present in the corresponding Pearson's χ^2 case (compare with the figure on page 5) have “gained weight”, and are visible here as short segments.

If we ask again to what value we must increase Δ to obtain a minimum coverage of 68.27%, this time the answer is $\Delta = 2.581$. This is also clearly worse than the corresponding Pearson's χ^2 case. The $\Delta = 2.581$ plot is shown here:

⁵Reference [7] shows that, when the variance is the comparison criterion, $-2\ln\lambda$ is superior to Pearson's χ^2 in the Poisson case.

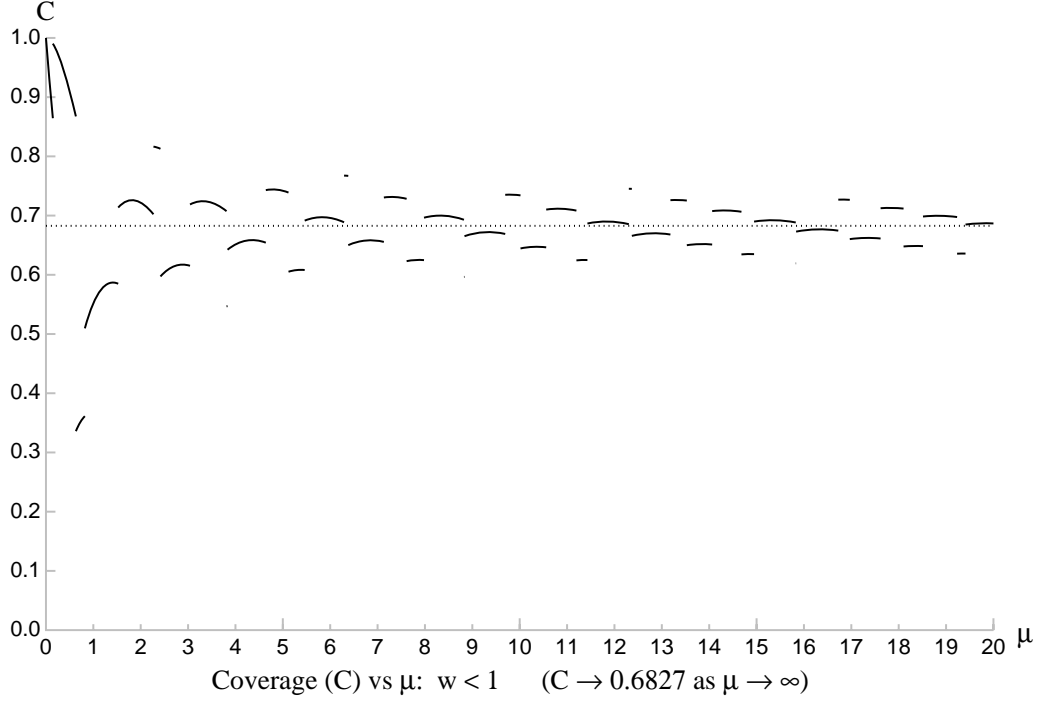


5 “Improved Likelihood Ratio” Intervals

The statistic

$$w = \frac{-2 \ln \lambda}{1 + \frac{1}{6}\mu^{-1}}$$

is the “improved likelihood ratio” [8, 9, 10] statistic for the Poisson case. As shown in [7], the mean and variance of the Poisson $-2 \ln \lambda$ are asymptotically $1 + \frac{1}{6}\mu^{-1} + O(\mu^{-2})$ and $2 + \frac{2}{3}\mu^{-1} + O(\mu^{-2})$ respectively, when expanded in powers of μ^{-1} . The rationale of the improved likelihood ratio is that w as defined above then has mean $1 + O(\mu^{-2})$ and variance $2 + O(\mu^{-2})$; i.e., closer to the moments of the χ^2 distribution for 1 degree of freedom (for large μ). The coverage of the resulting intervals is shown here:



The coverage looks qualitatively similar to that of the $-2 \ln \lambda$ case shown on page 10.

6 Classical-Frequentist Central-Intervals

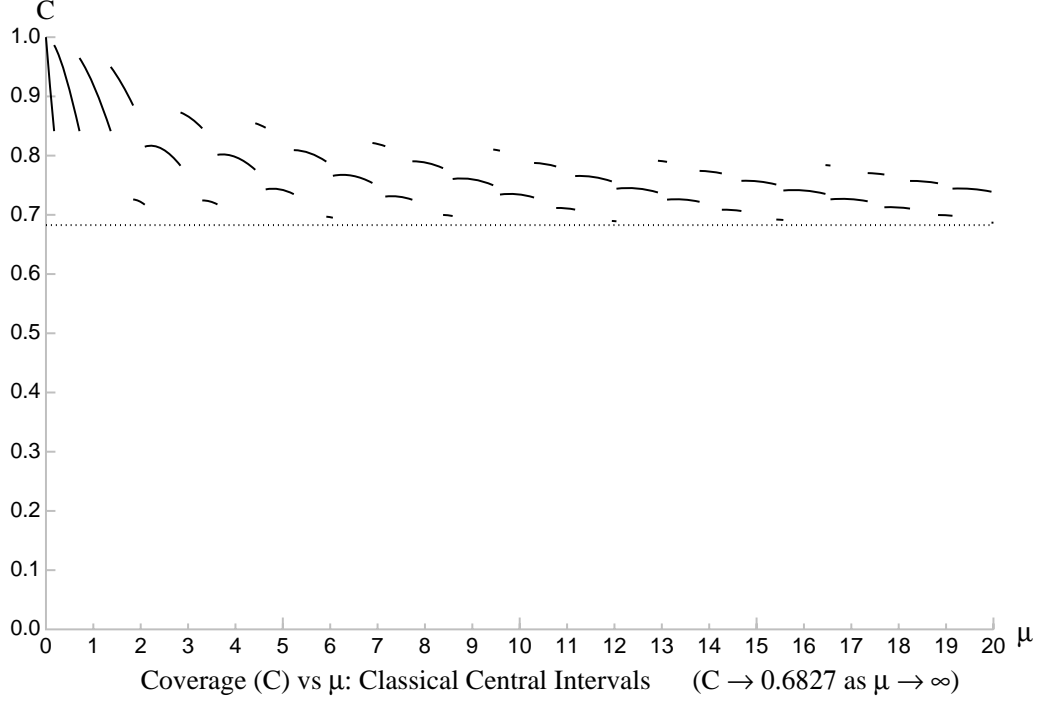
Since, from the frequentist point of view, none of the previous interval schemes have adequate coverage at small μ , we next investigate the coverage achieved by the “68.27%” (central) intervals of the classical frequentist approach. The classical approach to Poisson frequentist intervals is described in Ref. [11]. In this case, the (central) error interval for n observed events is given by the set of all μ such that:

$$\sum_{k=0}^n \frac{e^{-\mu} \mu^k}{k!} \geq \frac{1 - C_0}{2} \quad \text{and} \quad \sum_{k=n}^{\infty} \frac{e^{-\mu} \mu^k}{k!} \geq \frac{1 - C_0}{2}$$

The interval for $n = 0$ is defined completely by

$$\mu \leq \ln \frac{2}{1 - C_0}$$

since, when $n = 0$, the 2nd inequality, becoming $1 \geq \frac{1-C_0}{2}$, is true for all μ . The corresponding coverage plot is shown here:



The definition of these intervals is tailored so that the minimum coverage is guaranteed to be $\geq C_0$. The average overcoverage is worse here than for the case of the unified intervals (see the figure on page 19) considered next. This seems to be because conservatism is applied twice—there are two inequalities that both need to be satisfied—while the unified approach leads to only a single inequality.

7 Unified Intervals

We next investigate the coverage achieved by the “68.27%” intervals (zero background) of the unified approach[12]. As in the classical frequentist approach, the unified intervals will guarantee that the minimum coverage is $\geq C_0$. The error interval for n observed events is given by the set of all μ satisfying

$$U(\mu, n) = \sum_{k \in \mathcal{A}(\mu, n)} \frac{e^{-\mu} \mu^k}{k!} < C_0$$

where the set (of zero or more non negative integers) $\mathcal{A}(\mu, n)$ is defined as

$$\mathcal{A}(\mu, n) = \{ k \mid k \in \mathbf{Z}, k \geq 0, \text{ and } -2 \ln \lambda(\mu, k) < -2 \ln \lambda(\mu, n) \}$$

and \mathbf{Z} denotes the set of integers. The fact that the set $\mathcal{A}(\mu, n)$ is selected using the likelihood ratio is the hallmark of the unified approach. We might (crudely) describe $\mathcal{A}(\mu, n)$ as the set of all integers that give a “better fit” to μ than n does, where “better fit” is defined in terms of the likelihood ratio. Note that $n \notin \mathcal{A}(\mu, n)$.

Although the terse definition given above is mathematically equivalent to that of Ref. [12], this fact may not be obvious at first glance to readers familiar with that reference. For those readers, the brief explanation in the next paragraph will help to make the connection. (The rest, especially readers unfamiliar with the Neyman construction, may wish to skip the next paragraph entirely.)

In performing the Neyman construction for the Poisson case using likelihood ratio ordering, when constructing the band at any given fixed value of μ , one includes integers k in the band—starting with the one that gives the smallest $-2 \ln \lambda(\mu, k)$, and continuing, one by one, in order of smallest $-2 \ln \lambda(\mu, k)$ not yet included—only stopping when the probability (given the specified μ) of observing a k within the band finally becomes $\geq C_0$. When one reaches the stage of the construction at which the band is equal to the set $\mathcal{A}(\mu, n)$, the next integer to be considered for inclusion in the band is n , since, by our definition of $\mathcal{A}(\mu, n)$, all k with $-2 \ln \lambda(\mu, k) < -2 \ln \lambda(\mu, n)$ have already been included in the band. Then, if $\sum_{k \in \mathcal{A}(\mu, n)} e^{-\mu} \mu^k / k! \geq C_0$, the construction of the band (at this μ) has already terminated, meaning n will not be included in the band. On the other hand, if $\sum_{k \in \mathcal{A}(\mu, n)} e^{-\mu} \mu^k / k! < C_0$, the construction of the band must continue, meaning n will be included in the band—independent of what the value of $e^{-\mu} \mu^n / n!$ actually is, or how many additional integers k must be included in the band subsequent to the inclusion of n .

For example, suppose we observe n events, and we want to know if $\mu = n$ is within the error interval for that case. Since $\mathcal{A}(n, n) = \emptyset$ (empty set), $U(n, n) = 0$, which being less than C_0 , means that $\mu = n$ is included within the error interval for any $C_0 > 0$ and any n .

Another example: We observe $n = 1$ and we want to know if $\mu = 0.35$ is within the interval for $C_0 = 0.6827$. We have $\mathcal{A}(0.35, 1) = \{0\}$, and

therefore $U(0.35, 1) = e^{-0.35} = 0.7047$. Since $0.7047 \not\leq 0.6827$, $\mu = 0.35$ is not within the $n = 1$ interval. Based on this, one might falsely suspect that the $n = 1$ interval extends no further down than $-\ln(C_0) = 0.3817$. However, checking $\mu = 0.375$ for coverage when $n = 1$, we have $-2 \ln \lambda(0.375, 0) \not\leq -2 \ln \lambda(0.375, 1)$, $\mathcal{A}(0.375, 1) = \emptyset$, $U(0.375, 1) = 0$, and thus $\mu = 0.375$ is contained within the interval. Since $\mathcal{A}(\mu, 1)$ changes from $\{0\}$ to \emptyset at $\mu = e^{-1}$, the low end of the interval for $n = 1$ is given by $\mu_1 = e^{-1} = 0.3679$.

Interestingly, in the unified approach, the width of the error intervals no longer increases at a steady rate with n . This behavior is seen in the following table, which shows the endpoints of the intervals, and their widths, for $n = 0$ to 30:

| n | μ_1 | μ_2 | $\mu_2 - \mu_1$ | $-2 \ln \lambda(\mu_1, n)$ | $-2 \ln \lambda(\mu_2, n)$ |
|-----|---------|---------|-----------------|----------------------------|----------------------------|
| 0 | 0.0000 | 1.2904 | 1.2904 | 0.0000 | 2.5807 |
| 1 | 0.3679 | 2.7505 | 2.3827 | 0.7358 | 1.4775 |
| 2 | 0.7358 | 4.2504 | 3.5147 | 1.4715 | 1.4853 |
| 3 | 1.1036 | 5.3012 | 4.1976 | 2.2073 | 1.1865 |
| 4 | 2.3359 | 6.7764 | 4.4405 | 0.9750 | 1.3356 |
| 5 | 2.7505 | 7.8064 | 5.0559 | 1.4775 | 1.1577 |
| 6 | 3.8231 | 9.2783 | 5.4552 | 1.0546 | 1.3256 |
| 7 | 4.2504 | 10.3006 | 6.0502 | 1.4853 | 1.1931 |
| 8 | 5.3012 | 11.3187 | 6.0175 | 1.1865 | 1.0852 |
| 9 | 6.3342 | 12.7905 | 6.4562 | 0.9911 | 1.2544 |
| 10 | 6.7764 | 13.8060 | 7.0296 | 1.3356 | 1.1617 |
| 11 | 7.8064 | 14.8194 | 7.0130 | 1.1577 | 1.0819 |
| 12 | 8.8291 | 16.2920 | 7.4629 | 1.0227 | 1.2456 |
| 13 | 9.2783 | 17.3043 | 8.0259 | 1.3256 | 1.1724 |
| 14 | 10.3006 | 18.3152 | 8.0145 | 1.1931 | 1.1075 |
| 15 | 11.3187 | 19.3249 | 8.0061 | 1.0852 | 1.0495 |
| 16 | 12.3338 | 20.7991 | 8.4653 | 0.9955 | 1.2039 |
| 17 | 12.7905 | 21.8084 | 9.0180 | 1.2544 | 1.1480 |
| 18 | 13.8060 | 22.8169 | 9.0109 | 1.1617 | 1.0971 |
| 19 | 14.8194 | 23.8247 | 9.0053 | 1.0819 | 1.0506 |
| 20 | 15.8310 | 25.3003 | 9.4692 | 1.0125 | 1.1972 |
| 21 | 16.2920 | 26.3079 | 10.0159 | 1.2456 | 1.1512 |
| 22 | 17.3043 | 27.3150 | 10.0108 | 1.1724 | 1.1087 |
| 23 | 18.3152 | 28.3216 | 10.0065 | 1.1075 | 1.0692 |
| 24 | 19.3249 | 29.3278 | 10.0029 | 1.0495 | 1.0325 |
| 25 | 20.3336 | 30.8049 | 10.4712 | 0.9973 | 1.1699 |
| 26 | 20.7991 | 31.8110 | 11.0120 | 1.2039 | 1.1328 |
| 27 | 21.8084 | 32.8169 | 11.0084 | 1.1480 | 1.0980 |
| 28 | 22.8169 | 33.8223 | 11.0054 | 1.0971 | 1.0653 |
| 29 | 23.8247 | 34.8275 | 11.0027 | 1.0506 | 1.0345 |
| 30 | 24.8319 | 36.3057 | 11.4739 | 1.0079 | 1.1648 |

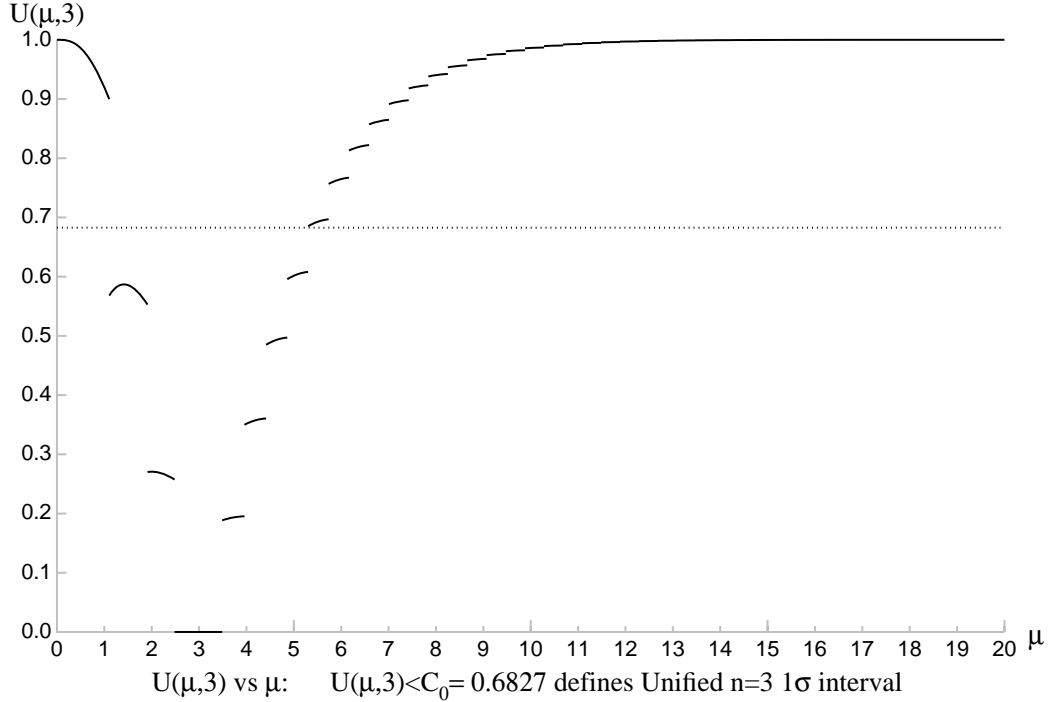
Unified 68.27% intervals

The width $\mu_2 - \mu_1$ of the intervals decreases slightly for the runs $n = 7$ to 8, $n = 10$ to 11, $n = 13$ to 15, $n = 17$ to 19, etc. It may seem strange that the error interval for $n = 8$ is slightly smaller than the error interval for

$n = 7$, for example. This, however, should not be taken too seriously. The interpretation of the size of an interval as proportional to the precision of the measurement is known to suffer when considerations of coverage become dominant. Some formulations of frequentist confidence intervals can occasionally even give empty confidence intervals, or intervals that just contain a single point. These are not to be interpreted as representing infinitely precise measurements: the coverage (in general) at any particular value of μ is not completely determined by the size and location of the interval for a single n , but has significant contributions from several overlapping intervals.

Comparing values from the μ_1 and μ_2 columns of the table, one notices that every value of μ_2 from 2.7505 to 23.8247 also appears somewhere in the μ_1 column. For $C_0 = 68.27\%$, this synchronization continues indefinitely when the table is extended. As observed in a previous section, this means that there are many orphan coverage points that will not show up in the coverage plot.

It is interesting to examine the calculation of the unified intervals in slightly more detail. We show $U(\mu, n)$ for the case $n = 3$ here:



As one would expect from its definition, $U(\mu, 3)$ as a function of μ has discontinuities at discrete points. As μ increases, the first discontinuity is lo-

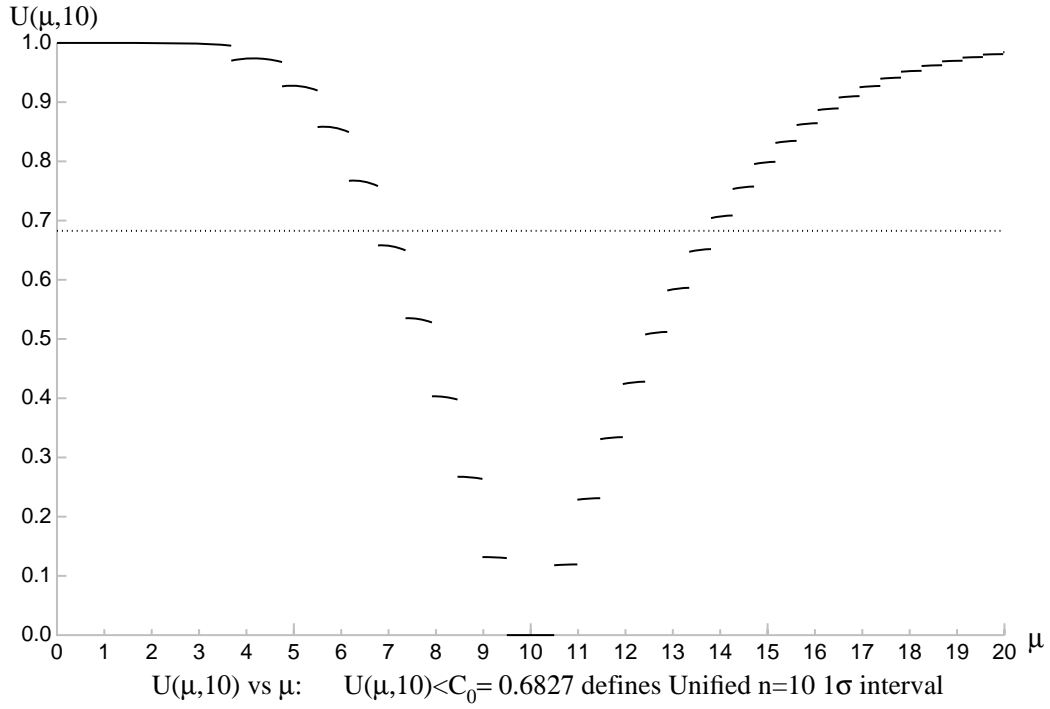
cated at $\mu = 3e^{-1} = 1.1036$, where $U(\mu, 3)$ drops below 68.27%, so we have $\mu_1 = 1.1036$. At $\mu = 5.3012$, $U(\mu, 3)$ jumps discontinuously from 0.6081 to 0.6852, so $\mu_2 = 5.3012$. Thus, the $n = 3$ interval's endpoints are defined by the location of the discontinuities in $U(\mu, 3)$.

Note that, if we tried to calculate the $n = 3$ interval for the case $C_0 = 57.5\%$, we would find that the equation $U(\mu, 3) = 0.575$ has two roots for $\mu < 3$, leading to an error-interval with a gap (or hole) in its interior, according to the definition given above. Error intervals with holes are generally considered unacceptable. The possibility of obtaining intervals with holes via the unified approach is noted in [12], which therefore adds an additional clause to the definition: any holes in the interior of an interval are to be added to the interval, so that the final interval can always be described simply as $[\mu_1, \mu_2]$.

It turns out that the hole-filling step never needs to be performed for the 68.27% (i.e. 1σ) intervals, but in general, one does need to check for this possibility. Surprisingly, for 1σ intervals, it seems that the equation $U(\mu, n) = C_0$ has no roots at all for $n > 0$. We conjecture⁶ that the endpoints of the 1σ intervals for $n > 0$ are always located at discontinuities in $U(\mu, n)$. This seems to be a property specific to the single point $C_0 = 0.682689492137\dots$ that is not shared by other values for C_0 in that neighborhood.

For general C_0 , what does happen is that the equation $U(\mu, n) = C_0$ cannot be satisfied for most (but not all) large n . This seems to be because the continuous segments of $U(\mu, n)$ flatten out as n becomes larger—illustrated here in the case $n = 10$:

⁶This conjecture is based on a numerical search covering $n = 1$ to $n = 10^7$. It would be nice to have a mathematical proof—or a counterexample.



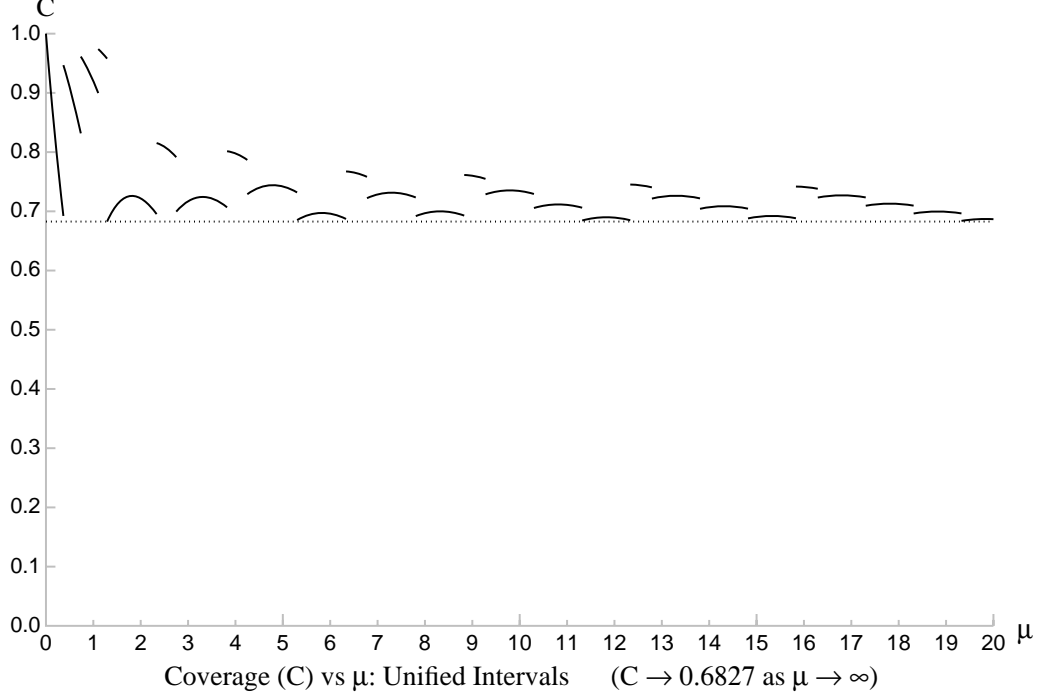
A random horizontal line drawn on this graph already has a relatively high probability of passing through a gap, rather than actually intersecting a segment.

The location of the discontinuities in $U(\mu, n)$ can be described analytically as follows: Solving $-2 \ln \lambda(\mu, k) = -2 \ln \lambda(\mu, n)$ for μ yields

$$\mu = n \exp \left(\frac{k \ln(n/k)}{n - k} - 1 \right) \quad (k \neq n)$$

Evaluating this expression for $k = 0, 1, 2, \dots, n-1, n+1, n+2, \dots$ therefore produces the location of the discontinuities of $U(\mu, n)$ in ascending order.

Finally, we show the coverage of the 1σ unified intervals:



As is desired, $C(\mu) \geq 68.27\%$ for all μ . If the previously mentioned conjecture is correct, $C(\mu) = C_0$ at only the single point $\mu \simeq 1.2904$, which is μ_2 for the $n = 0$ interval. At all other values of μ , $C(\mu) > C_0$. Because of the “synchronization” of the μ_1 and μ_2 mention above, this coverage plot looks qualitatively different from the previously shown coverage plots.

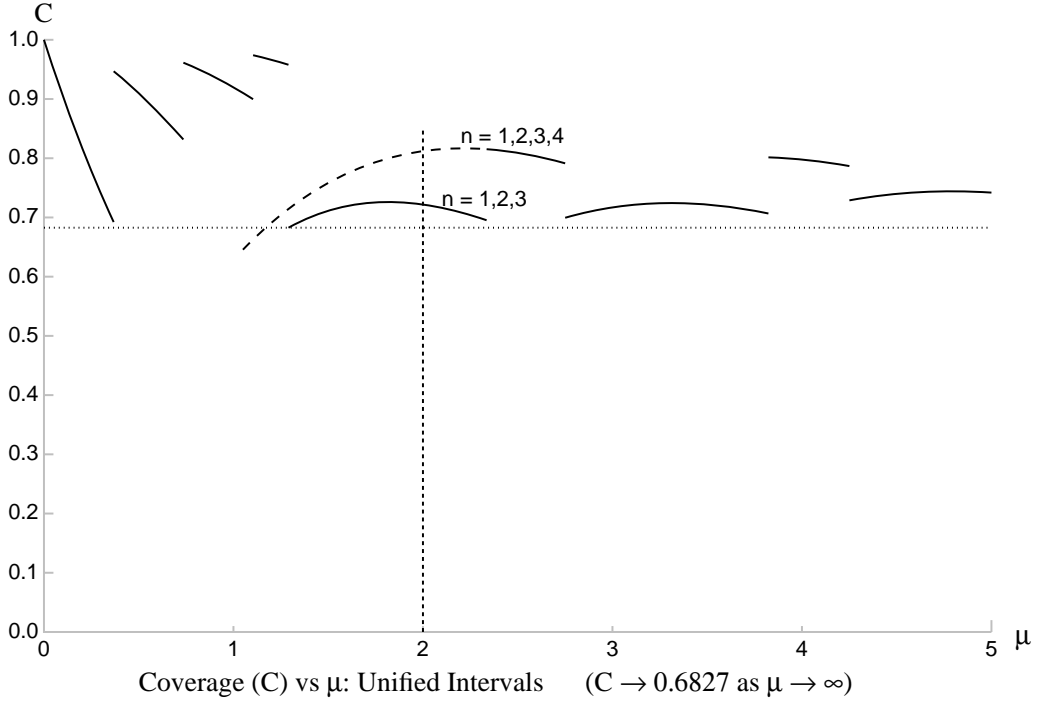
8 Interval Bias

There is an interesting concept that involves coverage considerations which can be introduced at this point: interval bias⁷. Quoting from reference [13]:

Further, it seems highly desirable that a good confidence interval should cover a value of θ with higher probability when it is the true value than when it is not, so that the confidence coefficient will exceed the probability of covering any false value. Such a confidence interval is called *unbiased*—this use of the term is unconnected with estimation bias.

⁷The concept of interval bias is directly related to test bias.

We can illustrate this concept using any of the coverage plots shown so far—they all demonstrate interval bias. Coverage is usually calculated for a given μ , assuming the “data” is generated with that μ . To investigate interval bias, we assume the data is generated at μ_{gen} , and the coverage is calculated for a different value μ_{cov} . For example, here we expand the μ -axis of the unified coverage plot from page 19, and extend the original coverage segment at $\mu \simeq 2.5$ as a dashed line down to lower values of μ :



As with the other coverage plots, the solid curves above give the coverage at the generated value (i.e. $\mu_{\text{gen}} = \mu_{\text{cov}}$). To determine the coverage for $\mu_{\text{cov}} = 2.5$ when the data is generated at $\mu_{\text{gen}} = 2.0$, the solid segment at $\mu = 2.5$ needs to be extended to the left, shown as dashed. Interval bias will be demonstrated by the fact that the dashed curve is higher than the solid curve at $\mu = 2$:

Suppose the true value of μ is 2.0. Then the coverage of that point under the unified scheme is

$$e^{-2} \left(2 + 2^2/2 + 2^3/6 \right) = \frac{16}{3} e^{-2} \simeq 0.7218$$

because the μ intervals for $n = 1, 2, 3$ include $\mu = 2$ (see Table on page 15). However, given that the true value of μ is 2.0, the coverage of the false value

$\mu = 2.5$ (or any false value in that segment) is

$$e^{-2} \left(2 + 2^2/2 + 2^3/6 + 2^4/24 \right) = 6e^{-2} \simeq 0.8120$$

because the μ intervals for $n = 1, 2, 3, 4$ include $\mu = 2.5$. That is, although the Poisson probabilities are calculated using the true value ($\mu = 2$) in both cases, the coverage probability of the false point $\mu = 2.5$ (or any false point in that segment) gains the $n = 4$ interval when compared to the coverage probability of the true value. Since this false value of the parameter μ is more likely to lie within the error bars than the true value, we have an example of interval bias.

It should be obvious both that interval bias is undesirable, and that is present in all the cases we have examined so far. The size of the bias is proportional to the size of the discontinuities in the $C(\mu)$ function, so we have another reason to want to keep the size of the “jumps” as small as possible. On the whole, because of the “synchronization” of the intervals, the unified approach does better than the others examined so far in keeping the size of the jumps small.

9 Pearson’s χ^2 Ordering

In analogy with the unified approach, we next consider intervals constructed by imposing an ordering based on Pearson’s χ^2 (instead of $-2 \ln \lambda$). Explicitly, the error interval for n observed events is given by the set of all μ satisfying

$$U(\mu, n) = \sum_{k \in \mathcal{A}(\mu, n)} \frac{e^{-\mu} \mu^k}{k!} < C_0$$

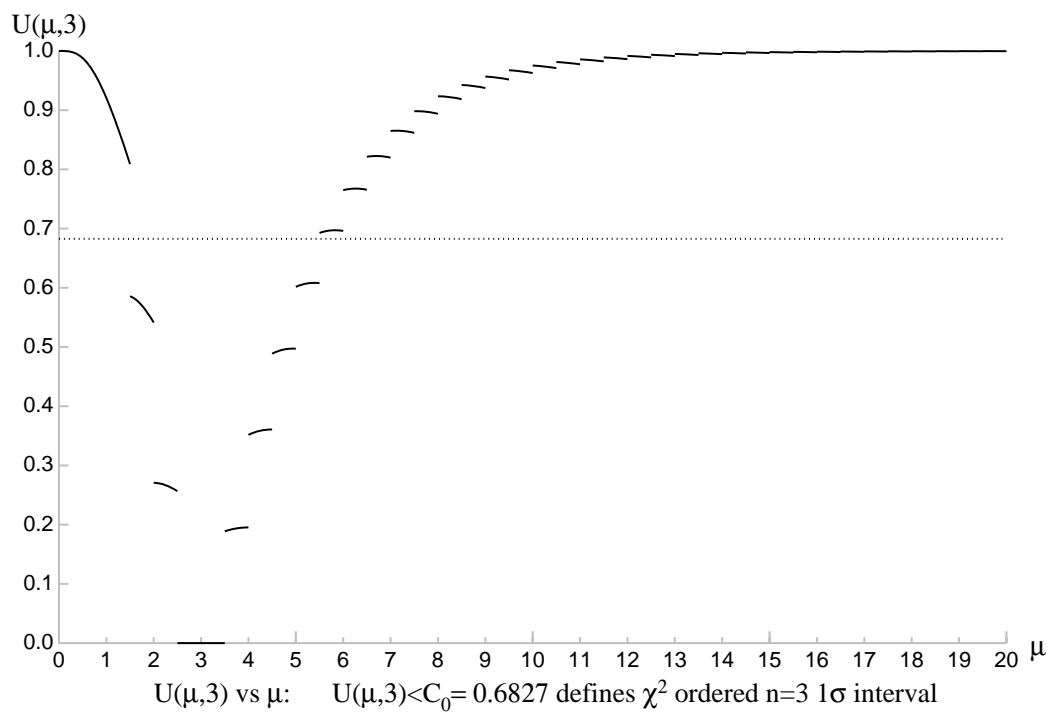
where now the set $\mathcal{A}(\mu, n)$ is defined as

$$\mathcal{A}(\mu, n) = \{ k \mid k \in \mathbf{Z}, k \geq 0, \text{ and } \frac{(k-\mu)^2}{\mu} < \frac{(n-\mu)^2}{\mu} \}$$

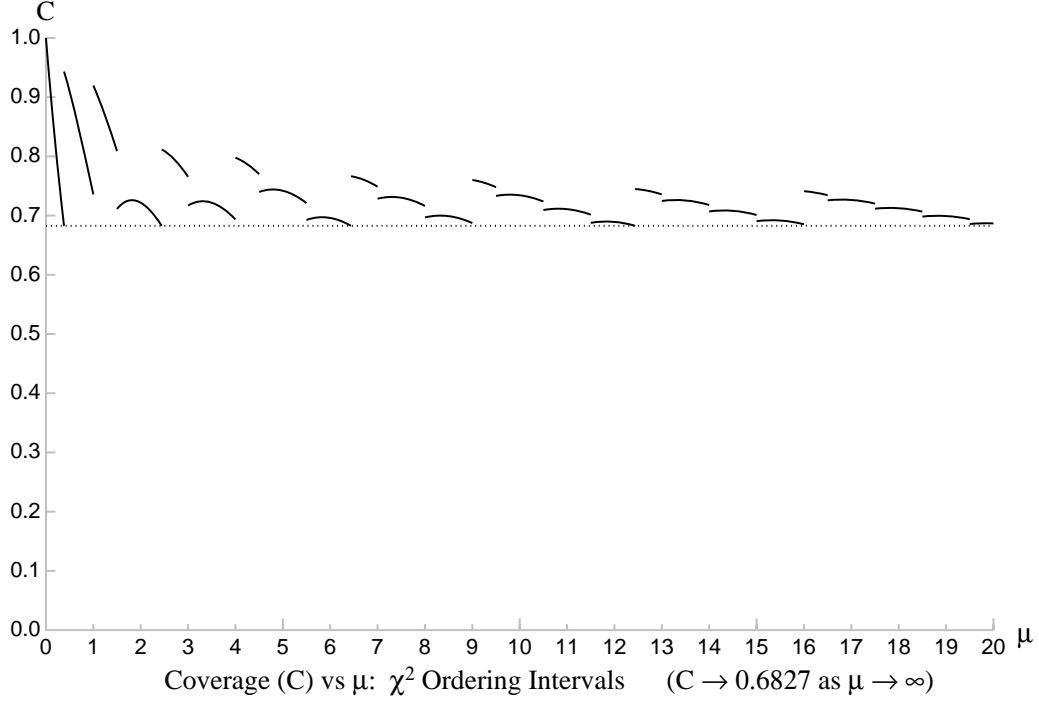
The locations of the discontinuities of $U(\mu, n)$, which occur at values of μ that satisfy $(k - \mu)^2/\mu = (n - \mu)^2/\mu$, are then simply given by

$$\mu = \frac{k + n}{2} \quad (k \neq n)$$

for $k = 0, 1, 2, \dots, n-1, n+1, n+2, \dots$. The function $U(\mu, n)$ for the case $n = 3$ is shown here:



The resulting coverage function $C(\mu)$ is shown here:



It is qualitatively similar to the unified coverage plot on page 19.

10 Probability Ordering

Instead of $-2 \ln \lambda$ or Pearson's χ^2 , we can also order on the Poisson probability itself. In this case, the error interval for n observed events is given by the set of all μ satisfying

$$U(\mu, n) = \sum_{k \in \mathcal{A}(\mu, n)} \frac{e^{-\mu} \mu^k}{k!} < C_0$$

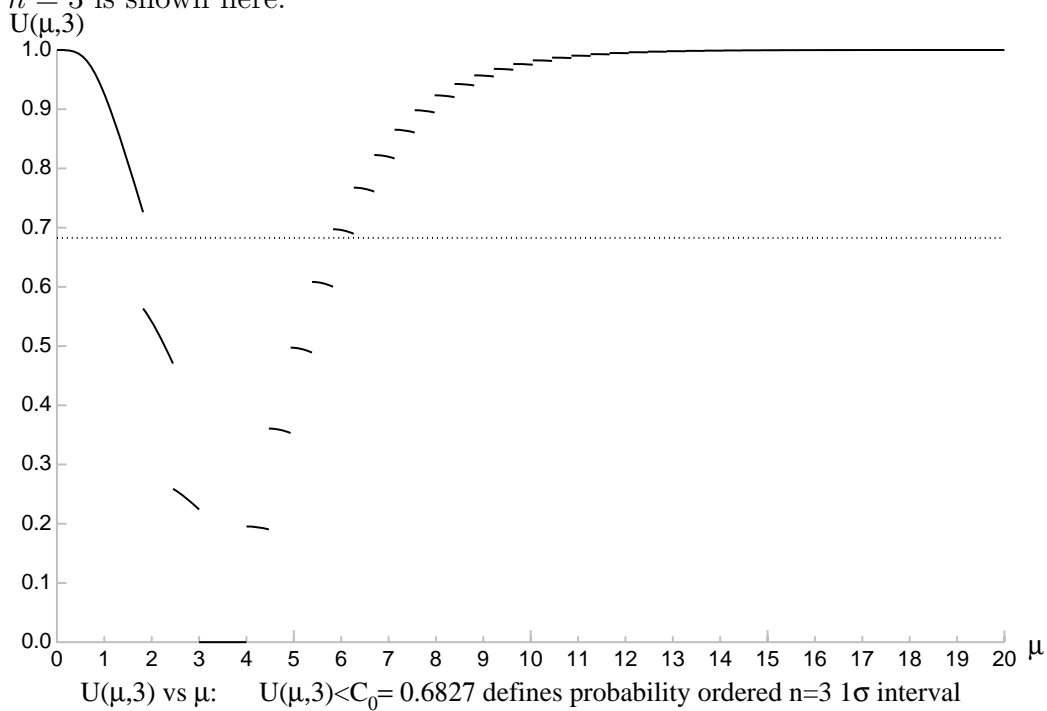
with the set $\mathcal{A}(\mu, n)$ defined as

$$\mathcal{A}(\mu, n) = \{ k \mid k \in \mathbf{Z}, k \geq 0, \text{ and } \frac{e^{-\mu} \mu^k}{k!} > \frac{e^{-\mu} \mu^n}{n!} \}$$

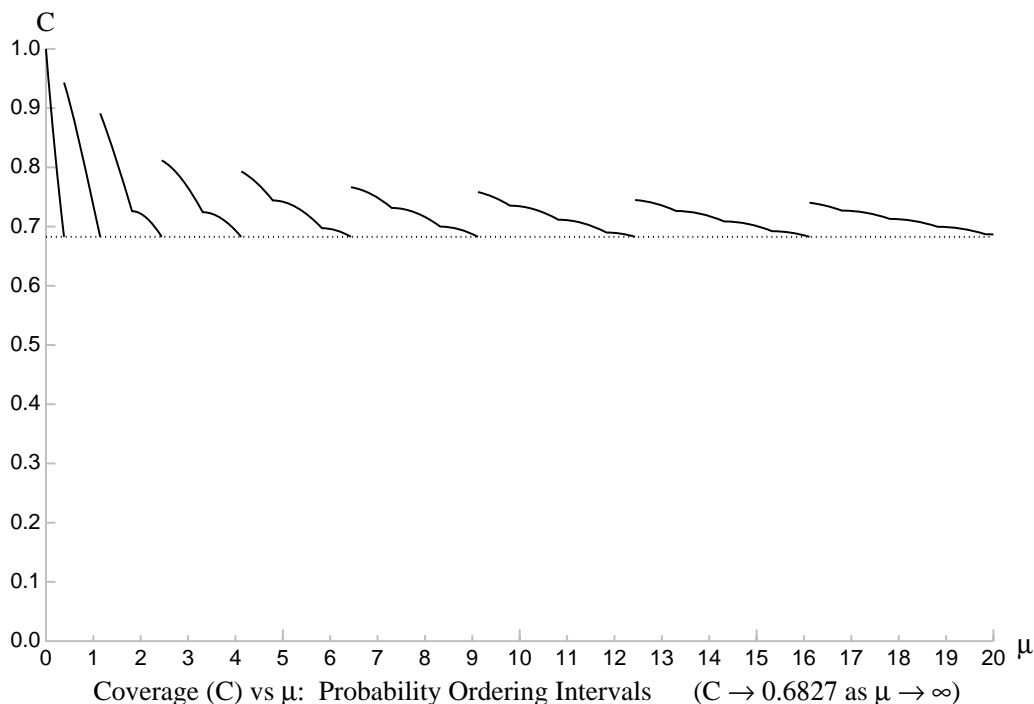
The locations of the discontinuities of $U(\mu, n)$ are then given by

$$\mu = \left(\frac{n!}{k!} \right)^{\frac{1}{n-k}} \quad (k \neq n)$$

for $k = 0, 1, 2, \dots, n-1, n+1, n+2, \dots$. The function $U(\mu, n)$ for the case $n = 3$ is shown here:



The resulting coverage function $C(\mu)$ is shown here:



Although quantitatively it is not so different from the unified coverage plot on page 19 or the Pearson's χ^2 ordering coverage plot on page 23, it seems at first glance quite different because the smaller discontinuities in $C(\mu)$ are “closed up”, leaving only a discontinuity in the first derivative⁸. At these points of discontinuity in the derivative, $C'(\mu)$ actually changes from a negative value to exactly zero. This behavior represents an improvement, since it eliminates the interval bias that was present in the neighborhood of the small discontinuities, but the main discontinuities still remain about the same.

11 Summary and Conclusions

- The intervals (or error bars) based on the change in the value of Pearson's χ^2 , Neyman's modified χ'^2 , the likelihood ratio ($-2 \ln \lambda$), and the “improved likelihood ratio” all produce both overcoverage and undercoverage (at different values of μ). Of these four schemes, the Pearson's χ^2 approach seems to give the best results. However, (any) undercoverage is deemed unacceptable by many frequentists—all of these methods

⁸We ignore the orphan coverage points present at these locations.

fail that test.

- Classical central intervals, unified intervals, Pearson's χ^2 ordered intervals, and probability ordered intervals all effectively eliminate undercoverage. The overcoverage of the classical central intervals is clearly worse than in the other three cases. The coverage functions $C(\mu)$ of the other three are quite similar, with probability ordered intervals arguably giving slightly better properties. However, all these conclusions only apply to the specific case investigated here: Poisson data with no background.
- Overcoverage is undesirable, so one is justified in trying to minimize it. Physicists intuitively expect exact coverage for error bars, but unfortunately, exact coverage is not attainable through normal means for discrete distributions—the Poisson case being a prime example. Interval bias (a wrong value of μ covering at a higher rate than the correct value of μ), like non exact coverage, is both undesirable and unavoidable in the Poisson case (and in the general discrete case).

Appendix: Interval Tables

| n | Pearson's χ^2 | | Neyman's χ'^2 | | Likelihood | | Improved L | |
|-----|--------------------|---------|--------------------|---------|------------|---------|------------|---------|
| | μ_1 | μ_2 | μ_1 | μ_2 | μ_1 | μ_2 | μ_1 | μ_2 |
| 0 | 0.0000 | 1.0000 | 0.0000 | 0.0000 | 0.0000 | 0.5000 | 0.0000 | 0.6319 |
| 1 | 0.3820 | 2.6180 | 0.0000 | 2.0000 | 0.3017 | 2.3577 | 0.1457 | 2.4170 |
| 2 | 1.0000 | 4.0000 | 0.5858 | 3.4142 | 0.8976 | 3.7654 | 0.8212 | 3.8117 |
| 3 | 1.6972 | 5.3028 | 1.2679 | 4.7321 | 1.5840 | 5.0802 | 1.5252 | 5.1198 |
| 4 | 2.4384 | 6.5616 | 2.0000 | 6.0000 | 2.3185 | 6.3463 | 2.2691 | 6.3815 |
| 5 | 3.2087 | 7.7913 | 2.7639 | 7.2361 | 3.0841 | 7.5811 | 3.0408 | 7.6131 |
| 6 | 4.0000 | 9.0000 | 3.5505 | 8.4495 | 3.8719 | 8.7936 | 3.8329 | 8.8232 |
| 7 | 4.8074 | 10.1926 | 4.3542 | 9.6458 | 4.6765 | 9.9891 | 4.6408 | 10.0168 |
| 8 | 5.6277 | 11.3723 | 5.1716 | 10.8284 | 5.4946 | 11.1711 | 5.4615 | 11.1973 |
| 9 | 6.4586 | 12.5414 | 6.0000 | 12.0000 | 6.3237 | 12.3422 | 6.2926 | 12.3670 |
| 10 | 7.2984 | 13.7016 | 6.8377 | 13.1623 | 7.1619 | 13.5040 | 7.1326 | 13.5277 |
| 11 | 8.1459 | 14.8541 | 7.6834 | 14.3166 | 8.0080 | 14.6580 | 7.9802 | 14.6807 |
| 12 | 9.0000 | 16.0000 | 8.5359 | 15.4641 | 8.8609 | 15.8052 | 8.8344 | 15.8270 |
| 13 | 9.8599 | 17.1401 | 9.3944 | 16.6056 | 9.7198 | 16.9463 | 9.6944 | 16.9674 |
| 14 | 10.7251 | 18.2749 | 10.2583 | 17.7417 | 10.5840 | 18.0822 | 10.5596 | 18.1025 |
| 15 | 11.5949 | 19.4051 | 11.1270 | 18.8730 | 11.4529 | 19.2132 | 11.4295 | 19.2330 |
| 16 | 12.4689 | 20.5311 | 12.0000 | 20.0000 | 12.3262 | 20.3401 | 12.3035 | 20.3592 |
| 17 | 13.3467 | 21.6533 | 12.8769 | 21.1231 | 13.2033 | 21.4630 | 13.1814 | 21.4816 |
| 18 | 14.2280 | 22.7720 | 13.7574 | 22.2426 | 14.0839 | 22.5823 | 14.0627 | 22.6005 |
| 19 | 15.1125 | 23.8875 | 14.6411 | 23.3589 | 14.9679 | 23.6984 | 14.9472 | 23.7161 |
| 20 | 16.0000 | 25.0000 | 15.5279 | 24.4721 | 15.8548 | 24.8115 | 15.8347 | 24.8288 |
| 21 | 16.8902 | 26.1098 | 16.4174 | 25.5826 | 16.7445 | 25.9218 | 16.7250 | 25.9387 |
| 22 | 17.7830 | 27.2170 | 17.3096 | 26.6904 | 17.6368 | 27.0295 | 17.6178 | 27.0460 |
| 23 | 18.6782 | 28.3218 | 18.2042 | 27.7958 | 18.5315 | 28.1348 | 18.5129 | 28.1510 |
| 24 | 19.5756 | 29.4244 | 19.1010 | 28.8990 | 19.4285 | 29.2378 | 19.4103 | 29.2537 |
| 25 | 20.4751 | 30.5249 | 20.0000 | 30.0000 | 20.3276 | 30.3387 | 20.3098 | 30.3543 |
| 26 | 21.3765 | 31.6235 | 20.9010 | 31.0990 | 21.2287 | 31.4377 | 21.2113 | 31.4530 |
| 27 | 22.2798 | 32.7202 | 21.8038 | 32.1962 | 22.1317 | 32.5347 | 22.1146 | 32.5497 |
| 28 | 23.1849 | 33.8151 | 22.7085 | 33.2915 | 23.0364 | 33.6300 | 23.0197 | 33.6447 |
| 29 | 24.0917 | 34.9083 | 23.6148 | 34.3852 | 23.9429 | 34.7235 | 23.9264 | 34.7381 |
| 30 | 25.0000 | 36.0000 | 24.5228 | 35.4772 | 24.8509 | 35.8155 | 24.8347 | 35.8298 |
| 31 | 25.9098 | 37.0902 | 25.4322 | 36.5678 | 25.7605 | 36.9060 | 25.7446 | 36.9201 |
| 32 | 26.8211 | 38.1789 | 26.3431 | 37.6569 | 26.6715 | 37.9950 | 26.6558 | 38.0089 |
| 33 | 27.7337 | 39.2663 | 27.2554 | 38.7446 | 27.5838 | 39.0826 | 27.5685 | 39.0963 |
| 34 | 28.6477 | 40.3523 | 28.1690 | 39.8310 | 28.4975 | 40.1689 | 28.4824 | 40.1824 |
| 35 | 29.5628 | 41.4372 | 29.0839 | 40.9161 | 29.4125 | 41.2540 | 29.3976 | 41.2673 |
| 36 | 30.4792 | 42.5208 | 30.0000 | 42.0000 | 30.3286 | 42.3379 | 30.3139 | 42.3510 |
| 37 | 31.3967 | 43.6033 | 30.9172 | 43.0828 | 31.2459 | 43.4206 | 31.2314 | 43.4335 |
| 38 | 32.3153 | 44.6847 | 31.8356 | 44.1644 | 32.1643 | 44.5022 | 32.1500 | 44.5150 |
| 39 | 33.2350 | 45.7650 | 32.7550 | 45.2450 | 33.0838 | 45.5827 | 33.0697 | 45.5953 |
| 40 | 34.1557 | 46.8443 | 33.6754 | 46.3246 | 34.0043 | 46.6622 | 33.9904 | 46.6747 |
| 41 | 35.0774 | 47.9226 | 34.5969 | 47.4031 | 34.9258 | 47.7407 | 34.9121 | 47.7531 |
| 42 | 36.0000 | 49.0000 | 35.5193 | 48.4807 | 35.8482 | 48.8183 | 35.8347 | 48.8305 |
| 43 | 36.9235 | 50.0765 | 36.4426 | 49.5574 | 36.7716 | 49.8949 | 36.7582 | 49.9070 |
| 44 | 37.8479 | 51.1521 | 37.3668 | 50.6332 | 37.6958 | 50.9707 | 37.6826 | 50.9826 |
| 45 | 38.7732 | 52.2268 | 38.2918 | 51.7082 | 38.6209 | 52.0456 | 38.6079 | 52.0574 |
| 46 | 39.6993 | 53.3007 | 39.2177 | 52.7823 | 39.5468 | 53.1197 | 39.5339 | 53.1314 |
| 47 | 40.6261 | 54.3739 | 40.1443 | 53.8557 | 40.4735 | 54.1930 | 40.4608 | 54.2045 |
| 48 | 41.5538 | 55.4462 | 41.0718 | 54.9282 | 41.4010 | 55.2655 | 41.3884 | 55.2769 |
| 49 | 42.4822 | 56.5178 | 42.0000 | 56.0000 | 42.3293 | 56.3372 | 42.3168 | 56.3486 |
| 50 | 43.4113 | 57.5887 | 42.9289 | 57.0711 | 43.2583 | 57.4083 | 43.2459 | 57.4195 |

Lower and upper limits of "68.27%" intervals.

| n | Classical | | Unified | | χ^2 Order | | Probability Order | |
|-----|-----------|---------|---------|---------|----------------|---------|-------------------|---------|
| | μ_1 | μ_2 | μ_1 | μ_2 | μ_1 | μ_2 | μ_1 | μ_2 |
| 0 | 0.0000 | 1.8410 | 0.0000 | 1.2904 | 0.0000 | 1.5000 | 0.0000 | 1.8171 |
| 1 | 0.1728 | 3.2995 | 0.3679 | 2.7505 | 0.3817 | 3.0000 | 0.3817 | 3.3098 |
| 2 | 0.7082 | 4.6379 | 0.7358 | 4.2504 | 1.0000 | 4.5000 | 1.1447 | 4.7894 |
| 3 | 1.3673 | 5.9182 | 1.1036 | 5.3012 | 1.5000 | 5.5000 | 1.8171 | 5.8274 |
| 4 | 2.0857 | 7.1628 | 2.3359 | 6.7764 | 2.4438 | 7.0000 | 2.4438 | 7.2989 |
| 5 | 2.8403 | 8.3825 | 2.7505 | 7.8064 | 3.0000 | 8.0000 | 3.3098 | 8.3239 |
| 6 | 3.6201 | 9.5836 | 3.8231 | 9.2783 | 4.0000 | 9.5000 | 4.1226 | 9.7947 |
| 7 | 4.4185 | 10.7703 | 4.2504 | 10.3006 | 4.5000 | 10.5000 | 4.7894 | 10.8143 |
| 8 | 5.2316 | 11.9451 | 5.3012 | 11.3187 | 5.5000 | 11.5000 | 5.8274 | 11.8303 |
| 9 | 6.0565 | 13.1102 | 6.3342 | 12.7905 | 6.4382 | 13.0000 | 6.4382 | 13.3019 |
| 10 | 6.8913 | 14.2669 | 6.7764 | 13.8060 | 7.0000 | 14.0000 | 7.2989 | 14.3160 |
| 11 | 7.7344 | 15.4165 | 7.8064 | 14.8194 | 8.0000 | 15.0000 | 8.3239 | 15.3282 |
| 12 | 8.5847 | 16.5598 | 8.8291 | 16.2920 | 9.0000 | 16.5000 | 9.1183 | 16.8010 |
| 13 | 9.4413 | 17.6976 | 9.2783 | 17.3043 | 9.5000 | 17.5000 | 9.7947 | 17.8123 |
| 14 | 10.3035 | 18.8304 | 10.3006 | 18.3152 | 10.5000 | 18.5000 | 10.8143 | 18.8225 |
| 15 | 11.1706 | 19.9587 | 11.3187 | 19.3249 | 11.5000 | 19.5000 | 11.8303 | 19.8315 |
| 16 | 12.0422 | 21.0831 | 12.3338 | 20.7991 | 12.4363 | 21.0000 | 12.4363 | 21.3059 |
| 17 | 12.9178 | 22.2037 | 12.7905 | 21.8084 | 13.0000 | 22.0000 | 13.3019 | 22.3147 |
| 18 | 13.7971 | 23.3210 | 13.8060 | 22.8169 | 14.0000 | 23.0000 | 14.3160 | 23.3228 |
| 19 | 14.6798 | 24.4352 | 14.8194 | 23.8247 | 15.0000 | 24.0000 | 15.3282 | 24.3301 |
| 20 | 15.5656 | 25.5465 | 15.8310 | 25.3003 | 16.0000 | 25.5000 | 16.1169 | 25.8058 |
| 21 | 16.4542 | 26.6552 | 16.2920 | 26.3079 | 16.5000 | 26.5000 | 16.8010 | 26.8132 |
| 22 | 17.3455 | 27.7614 | 17.3043 | 27.3150 | 17.5000 | 27.5000 | 17.8123 | 27.8199 |
| 23 | 18.2393 | 28.8652 | 18.3152 | 28.3216 | 18.5000 | 28.5000 | 18.8225 | 28.8263 |
| 24 | 19.1354 | 29.9669 | 19.3249 | 29.3278 | 19.5000 | 29.5000 | 19.8315 | 29.8321 |
| 25 | 20.0337 | 31.0666 | 20.3336 | 30.8049 | 20.4355 | 31.0000 | 20.4355 | 31.3094 |
| 26 | 20.9340 | 32.1643 | 20.7991 | 31.8110 | 21.0000 | 32.0000 | 21.3059 | 32.3153 |
| 27 | 21.8362 | 33.2602 | 21.8084 | 32.8169 | 22.0000 | 33.0000 | 22.3147 | 33.3209 |
| 28 | 22.7403 | 34.3544 | 22.8169 | 33.8223 | 23.0000 | 34.0000 | 23.3228 | 34.3262 |
| 29 | 23.6461 | 35.4470 | 23.8247 | 34.8275 | 24.0000 | 35.0000 | 24.3301 | 35.3311 |
| 30 | 24.5535 | 36.5380 | 24.8319 | 36.3057 | 25.0000 | 36.5000 | 25.1162 | 36.8095 |
| 31 | 25.4624 | 37.6276 | 25.3003 | 37.3110 | 25.5000 | 37.5000 | 25.8058 | 37.8146 |
| 32 | 26.3729 | 38.7158 | 26.3079 | 38.3160 | 26.5000 | 38.5000 | 26.8132 | 38.8194 |
| 33 | 27.2847 | 39.8026 | 27.3150 | 39.3207 | 27.5000 | 39.5000 | 27.8199 | 39.8240 |
| 34 | 28.1979 | 40.8881 | 28.3216 | 40.3251 | 28.5000 | 40.5000 | 28.8263 | 40.8283 |
| 35 | 29.1123 | 41.9724 | 29.3278 | 41.3294 | 29.5000 | 41.5000 | 29.8321 | 41.8325 |
| 36 | 30.0280 | 43.0555 | 30.3335 | 42.8090 | 30.4351 | 43.0000 | 30.4351 | 43.3122 |
| 37 | 30.9449 | 44.1376 | 30.8049 | 43.8134 | 31.0000 | 44.0000 | 31.3094 | 44.3164 |
| 38 | 31.8628 | 45.2185 | 31.8110 | 44.8176 | 32.0000 | 45.0000 | 32.3153 | 45.3205 |
| 39 | 32.7819 | 46.2984 | 32.8169 | 45.8216 | 33.0000 | 46.0000 | 33.3209 | 46.3244 |
| 40 | 33.7020 | 47.3773 | 33.8223 | 46.8254 | 34.0000 | 47.0000 | 34.3262 | 47.3281 |
| 41 | 34.6231 | 48.4552 | 34.8275 | 47.8291 | 35.0000 | 48.0000 | 35.3311 | 48.3317 |
| 42 | 35.5452 | 49.5322 | 35.8323 | 49.3097 | 36.0000 | 49.5000 | 36.1159 | 49.8124 |
| 43 | 36.4682 | 50.6083 | 36.3057 | 50.3135 | 36.5000 | 50.5000 | 36.8095 | 50.8161 |
| 44 | 37.3921 | 51.6835 | 37.3110 | 51.3171 | 37.5000 | 51.5000 | 37.8146 | 51.8197 |
| 45 | 38.3168 | 52.7579 | 38.3160 | 52.3206 | 38.5000 | 52.5000 | 38.8194 | 52.8231 |
| 46 | 39.2424 | 53.8315 | 39.3207 | 53.3240 | 39.5000 | 53.5000 | 39.8240 | 53.8264 |
| 47 | 40.1688 | 54.9043 | 40.3251 | 54.3273 | 40.5000 | 54.5000 | 40.8283 | 54.8296 |
| 48 | 41.0960 | 55.9763 | 41.3294 | 55.3304 | 41.5000 | 55.5000 | 41.8325 | 55.8327 |
| 49 | 42.0240 | 57.0476 | 42.3335 | 56.8121 | 42.4349 | 57.0000 | 42.4349 | 57.3144 |
| 50 | 42.9527 | 58.1182 | 42.8090 | 57.8153 | 43.0000 | 58.0000 | 43.3122 | 58.3176 |

Lower and upper limits of “68.27%” intervals.

References

- [1] W.T. Eadie, D. Drijard, F.E. James, M. Roos, and B. Sadoulet “Statistical Methods in Experimental Physics”, (North-Holland Publishing Co, Amsterdam, 1971), §11.2, p 257.
- [2] Glen Cowan, “Statistical Data Analysis”, (Oxford University Press, Oxford, 1998), §4.7 p 61.
- [3] F. James, “Minuit, Function Minimization and Error Analysis”, CERN long writeup D506, wwwinfo.cern.ch/asdoc/minuit/minmain.html
- [4] Paul Harrison, “Blind Analyses” in Proceedings of the Conference on Advanced Techniques in Particle Physics, M. Whalley and L. Lyons (ed.), IPPP/02/39 (July 2002), page 278.
www.ippp.dur.ac.uk/Workshops/02/statistics/proceedings/harrison.ps
- [5] Harold Jeffreys, “Theory of Probability”, 3rd ed., (Oxford University Press, Oxford, 1961), §4.1, p 197; W.T. Eadie, *et al.*, *ibid.*, §8.4.5, p 171; Glen Cowan, *ibid.*, §7.4 p 101.
- [6] K. Hagiwara et al., Physical Review D **66**, 010001-229 (2002), Statistics Review, revised October 2001 by G. Cowan,
pdg.lbl.gov/2002/contents_sports.html#mathtoolsetc
- [7] J.G. Heinrich, “The Log Likelihood Ratio of the Poisson Distribution for Small μ ”, CDF note 5718, version 2,
www-cdf.fnal.gov/physics/statistics/notes/cdf5718_loglikeratv2.ps
- [8] W.T. Eadie, et al., *ibid.*, §10.5.5, p 236;
- [9] Alan Stuart, J. Keith Ord, and Steven Arnold, “Kendall’s Advanced Theory of Statistics”, Volume 2A 6th ed., (Oxford University Press, Oxford, 1999), §22.9, p 249.
- [10] Morten Frydenberg and Jens Ledet Jensen, “Is the ‘improved likelihood ratio statistic’ really improved in the discrete case?”, Biometrika **76** p 655 (1989).
- [11] Glen Cowan, *ibid.*, §9.4 p 126.

- [12] G.J. Feldman and R.D. Cousins, “Unified approach to the classical statistical analysis of small signals”, Physical Review D **57**, 3873 (1998).
link.aps.org/abstract/PRD/v57/p3873
www.hepl.harvard.edu/~feldman/Unified_Approach.ps
- [13] Alan Stuart et al., *ibid.*, §19.14 p 128.